

October 7 - 9
Groningen, The Netherlands
www.informationuniverse.rug.nl

THE INFORMATION UNIVERSE

What is the role of information
in the physics of our Universe?

Keynote speakers

Erik Verlinde (Univ. of Amsterdam)
Alex Szalay (Johns Hopkins Univ.)
Gerard 't Hooft (Univ. of Utrecht)
Gregory Chaitin (Univ. of Rio de Janeiro)
Charley Lineweaver (Australian National Univ.)
Lude Franke (Univ. of Groningen)

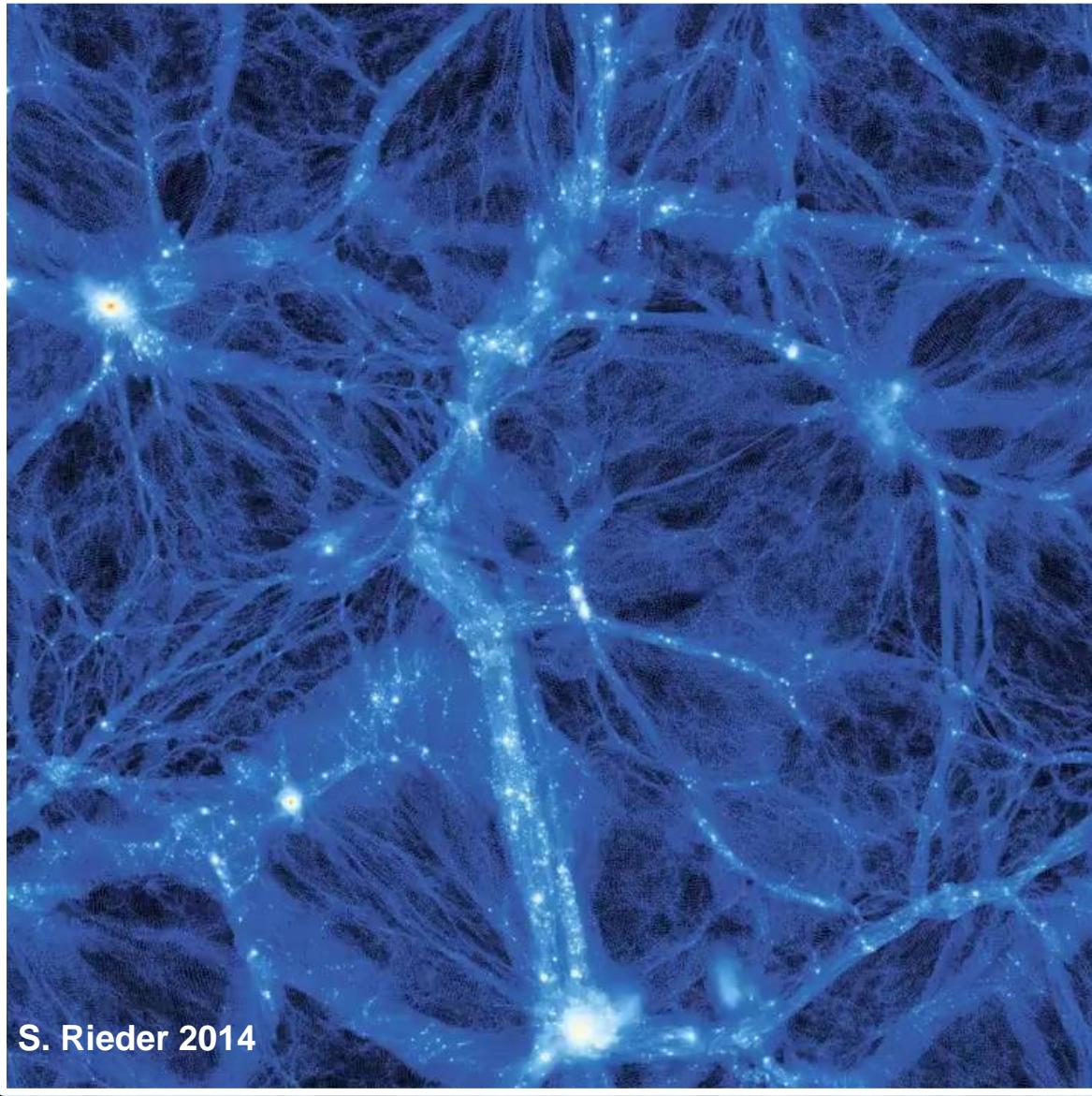
Scientific Organizing Committee

Edwin Valentijn (OmegaCEN)
Eric Bergshoeff (Van Swinderen Inst.)
Gert Vegter (Johan Bernoulli Inst.)
Rien van de Weijgaert (Kapteyn Astronomical Inst.)

het Kosmische Web: de Grootste Structuur in het Heelal

Rien van de Weijgaert, Inforversum, Groningen, 25 juni 2015





S. Rieder 2014

Cosmic Web

on scales of ~0.1-500
millions of lightyears

complex weblike pattern

in which
matter, gas & galaxies
are organized in

- ⌚ compact clusters,
- ⌚ elongated filaments
- ⌚ flattened walls
- ⌚ around
- ⌚ cosmic voids

Cosmic Fossil:

- our origin -

how did structure in the Universe emerge ?

Universe after 379000 years:

almost without
any profile

$$\frac{\Delta T}{T} < 10^{-5}$$

$$\frac{\Delta r}{r} \leq 1.4 \times 10^{-3}$$

$$\frac{\Delta r}{r} = 10^{-5} : r = 60.4 \text{ m}$$



Cosmic Dilemma:

379000 jaar after Big Bang:
uniform, nearly without any structure

How did the wealth of
objects and structure emerge?

Atlas of the Universe

Map of the Universe

How to map structures
in the Universe ?

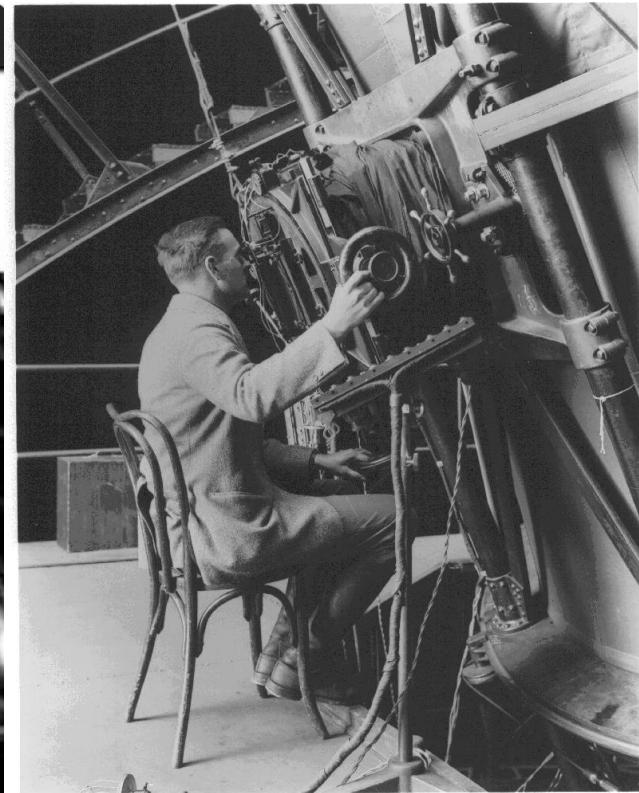
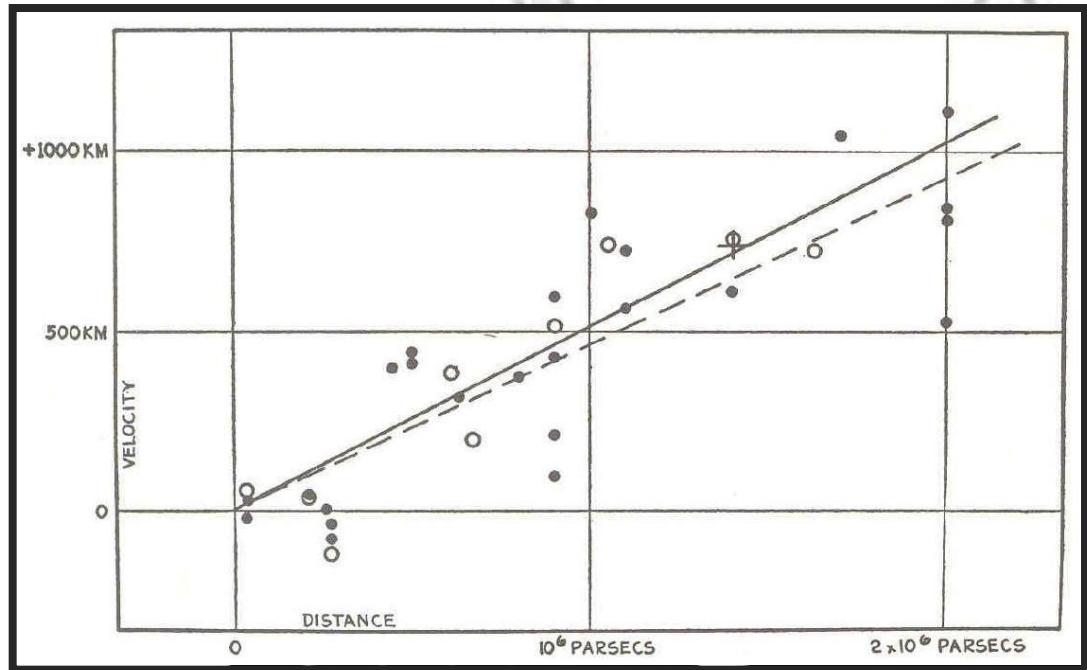
- Galaxies as light beacons
- Map of galaxy locations
- tracing of structures from distribution of galaxies



A dense field of galaxies in deep space, showing various colors and sizes of galactic clusters.

... Galaxies ...

Hubble Expansion



Distances in the Universe: redshift



moving toward you: blueshift



at rest

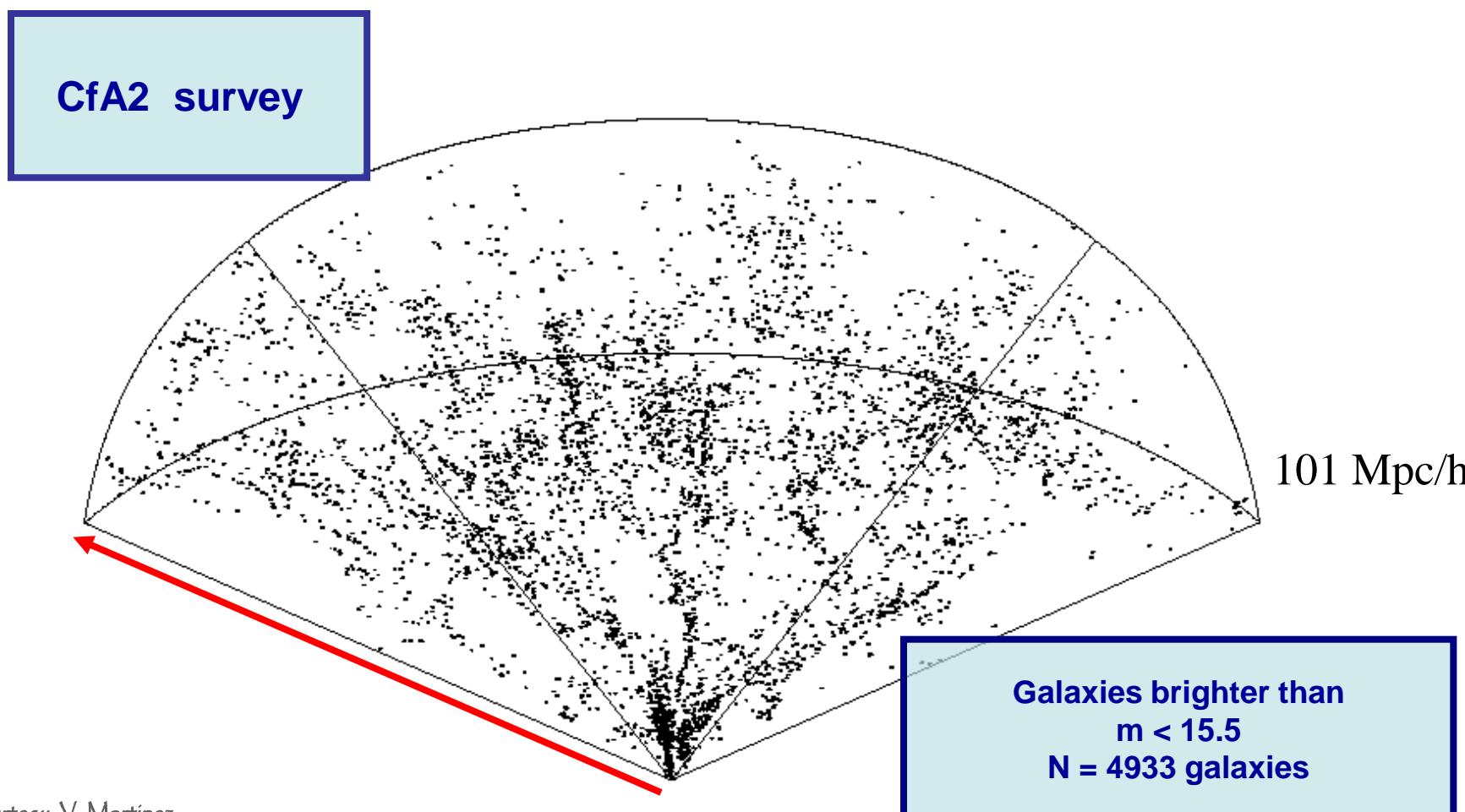
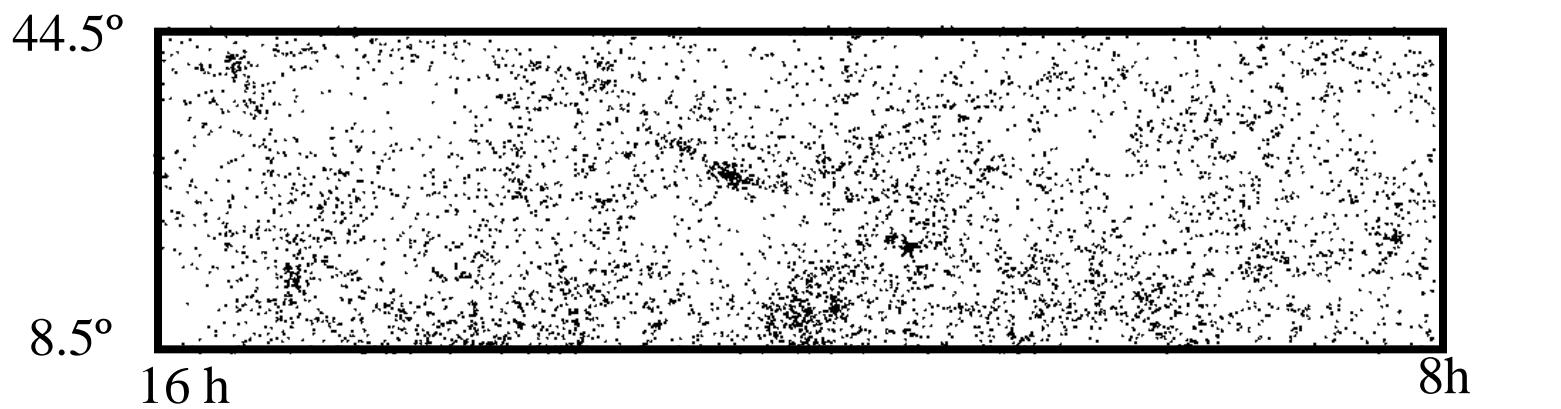


moving away from you: redshift

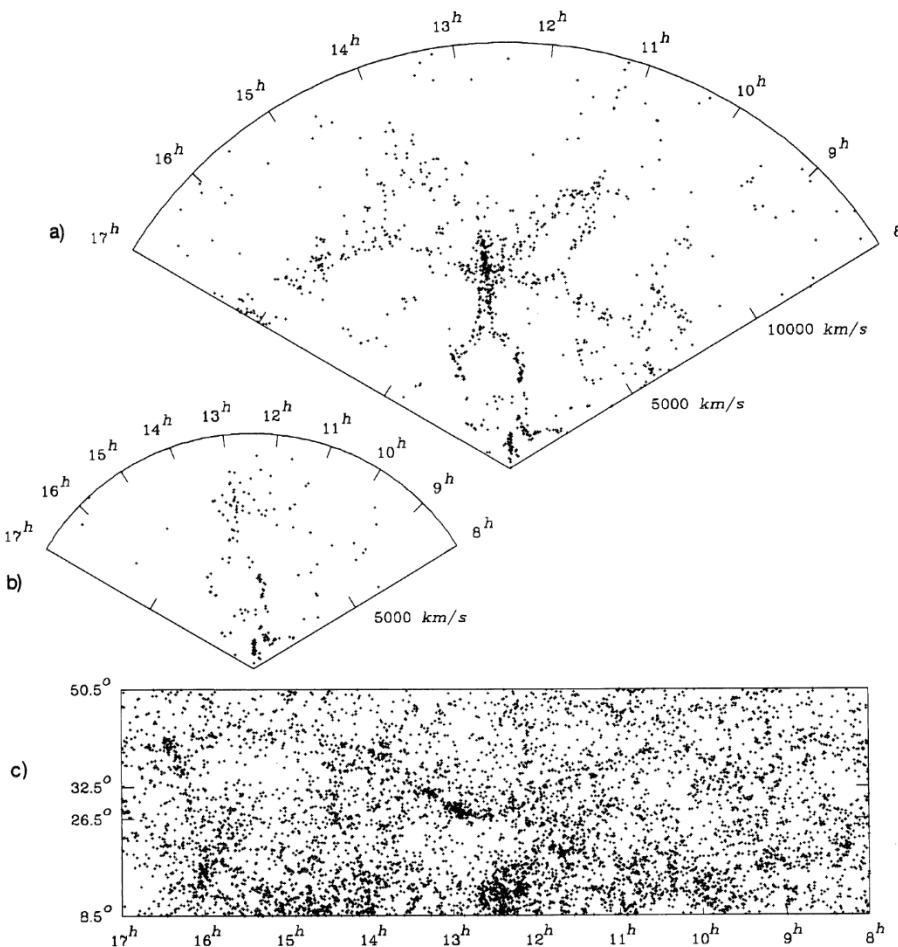


redshift ★ expansion velocity ★ Hubble: distance

$$cz = H r$$



“Stickman” & Soapsud

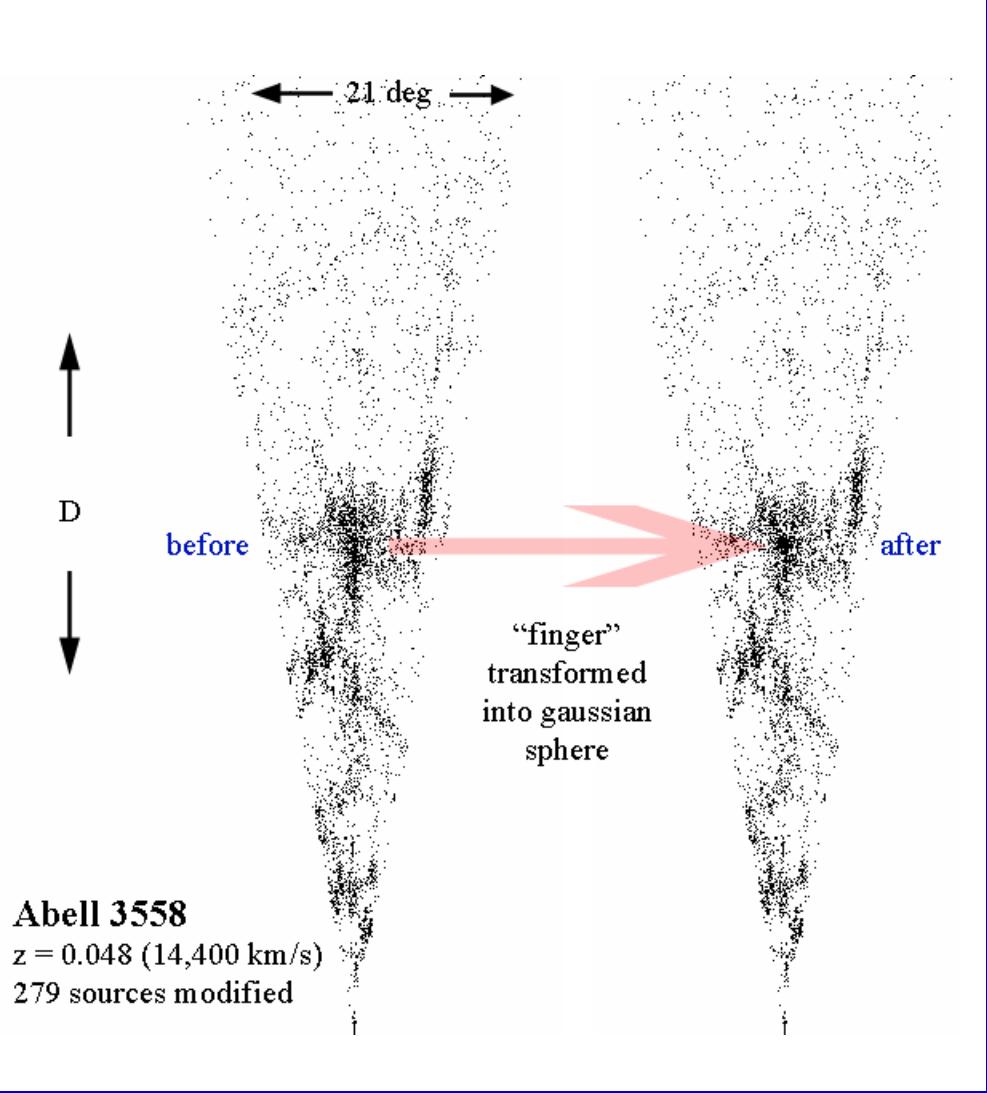
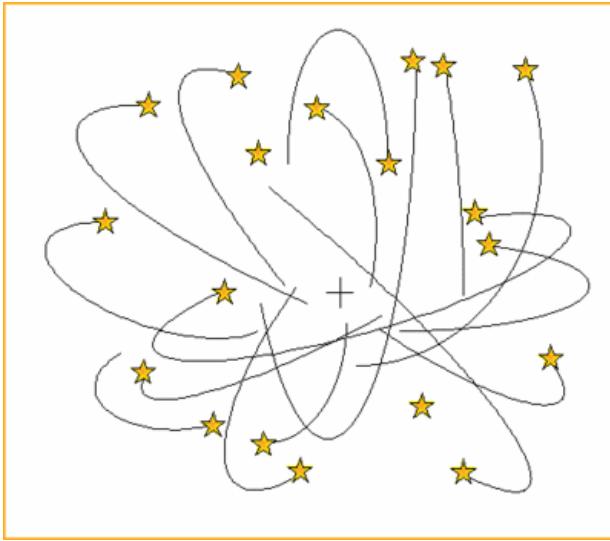


deLapparent, Geller & Huchra, 1986:

“a slice of the Universe”

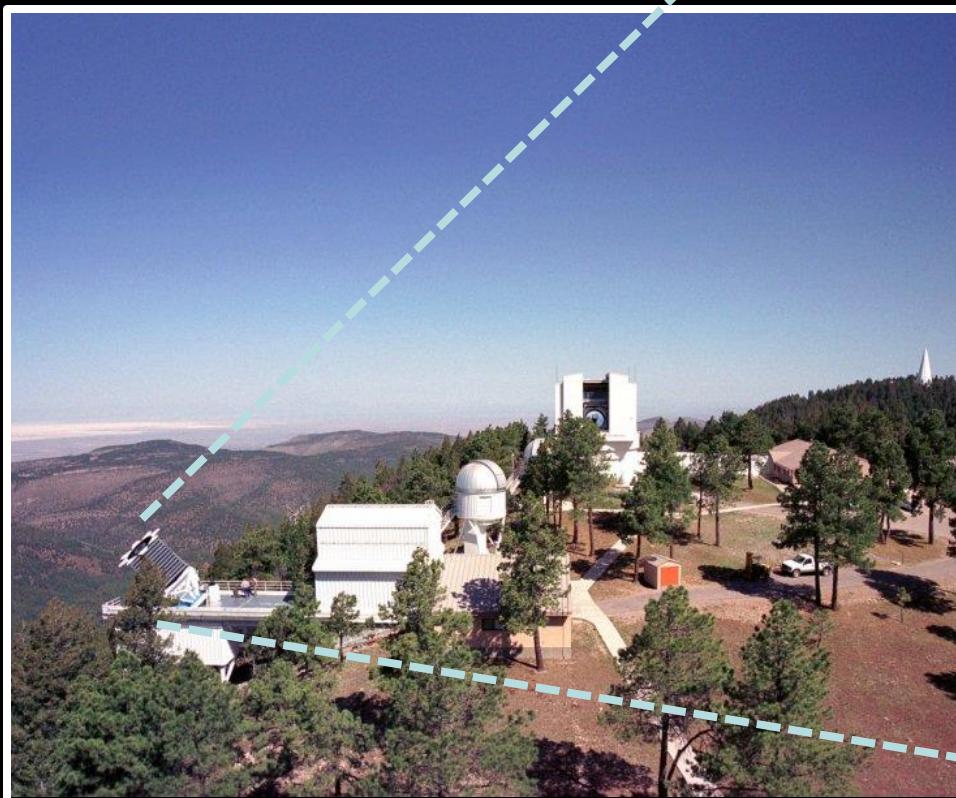
Voids represent a key element
of a galaxy distribution resembling
a soapsud.

Fingers of God



Sloan Digital Sky Survey (SDSS)

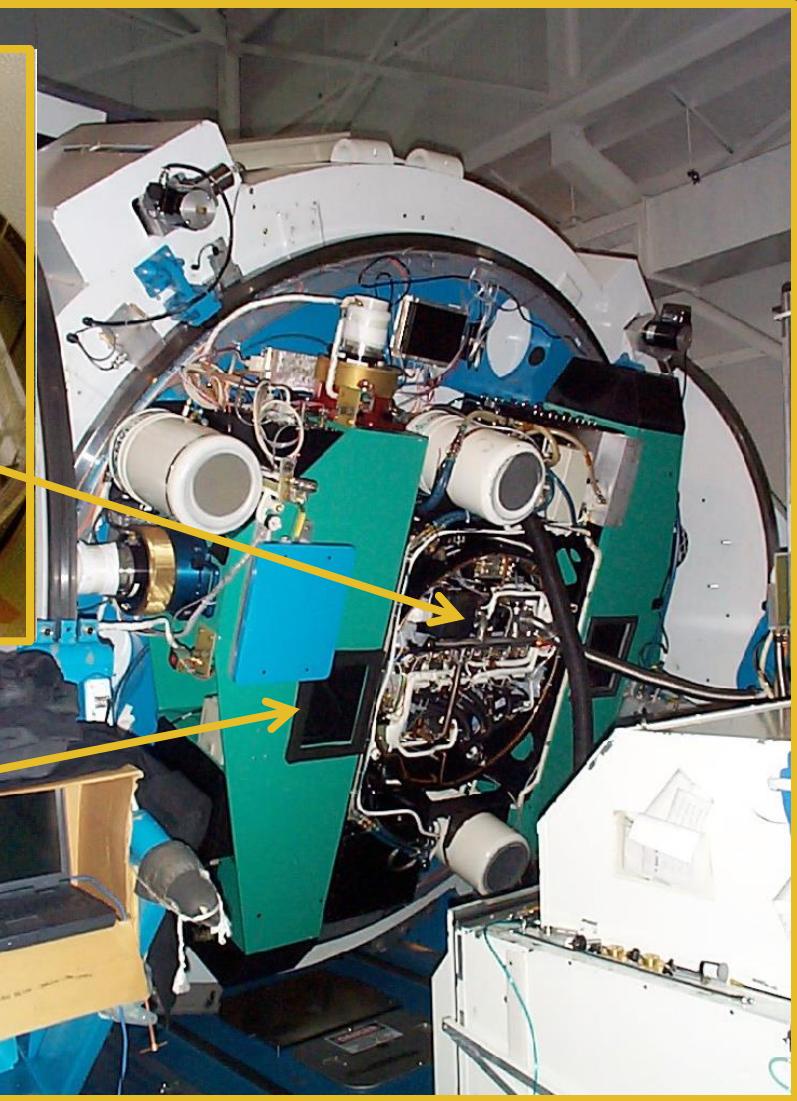
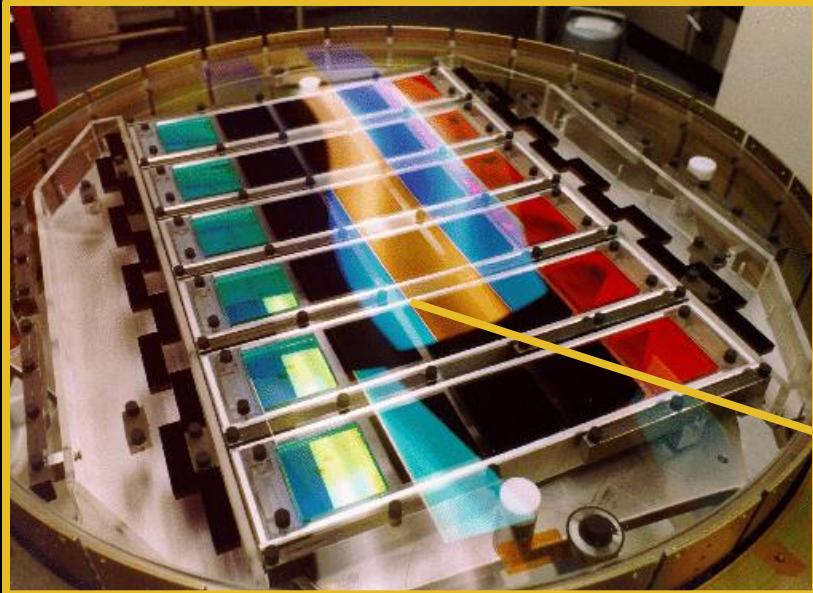
**special-purpose
2.5m wide-angle telescope
Apache Point Observatory (New Mexico)**



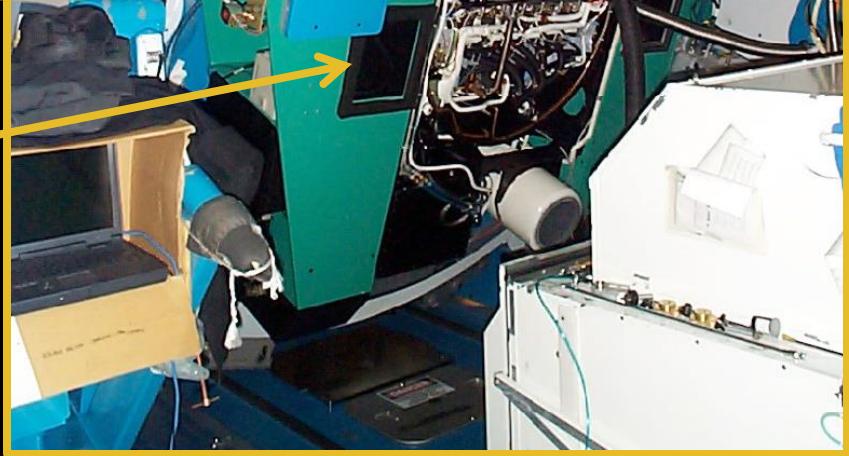
SDSS survey

5-color
Camera

30 CCD
chips



Fiber
Spectrograph

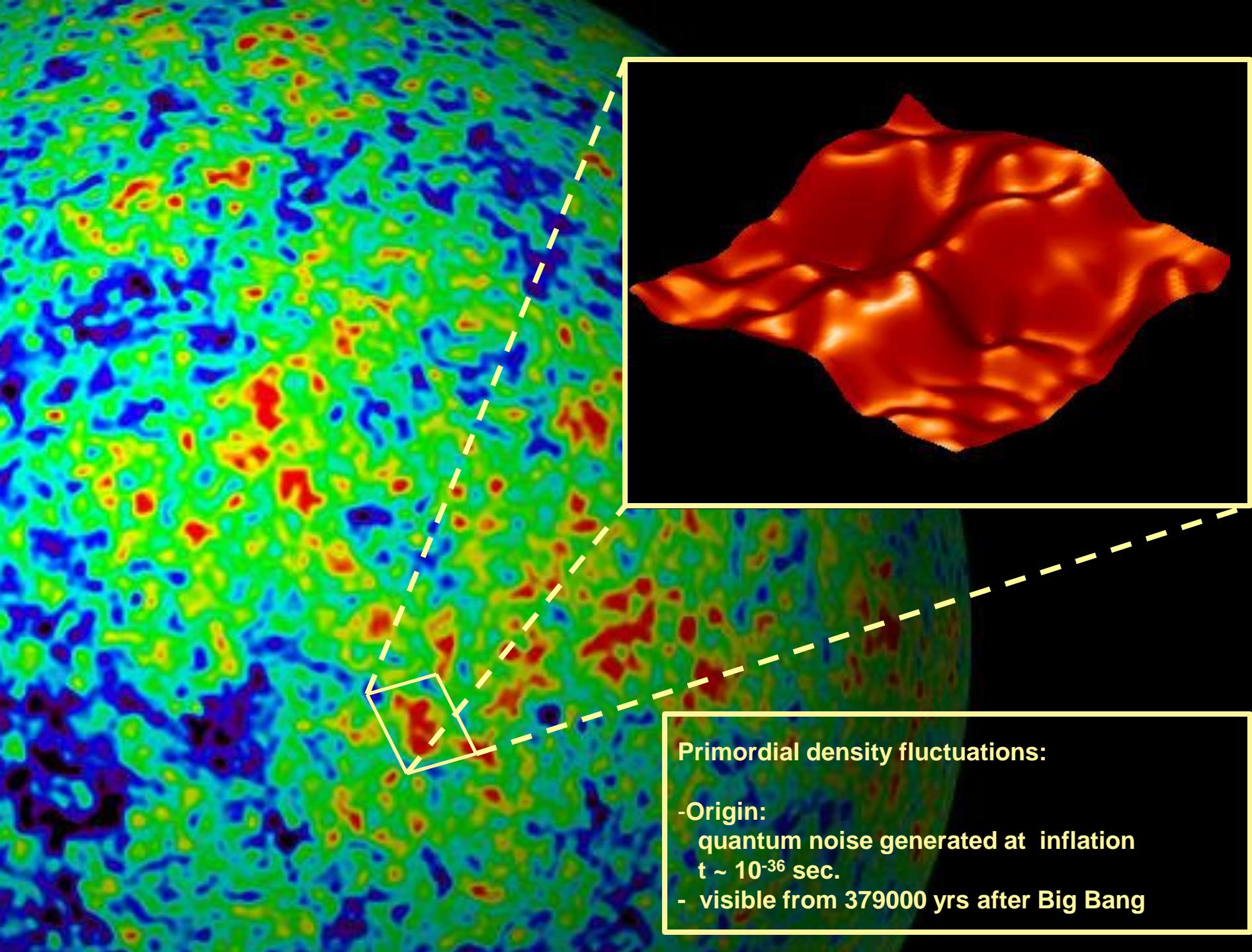


full-dome: SDSS-1

Structure Formation:

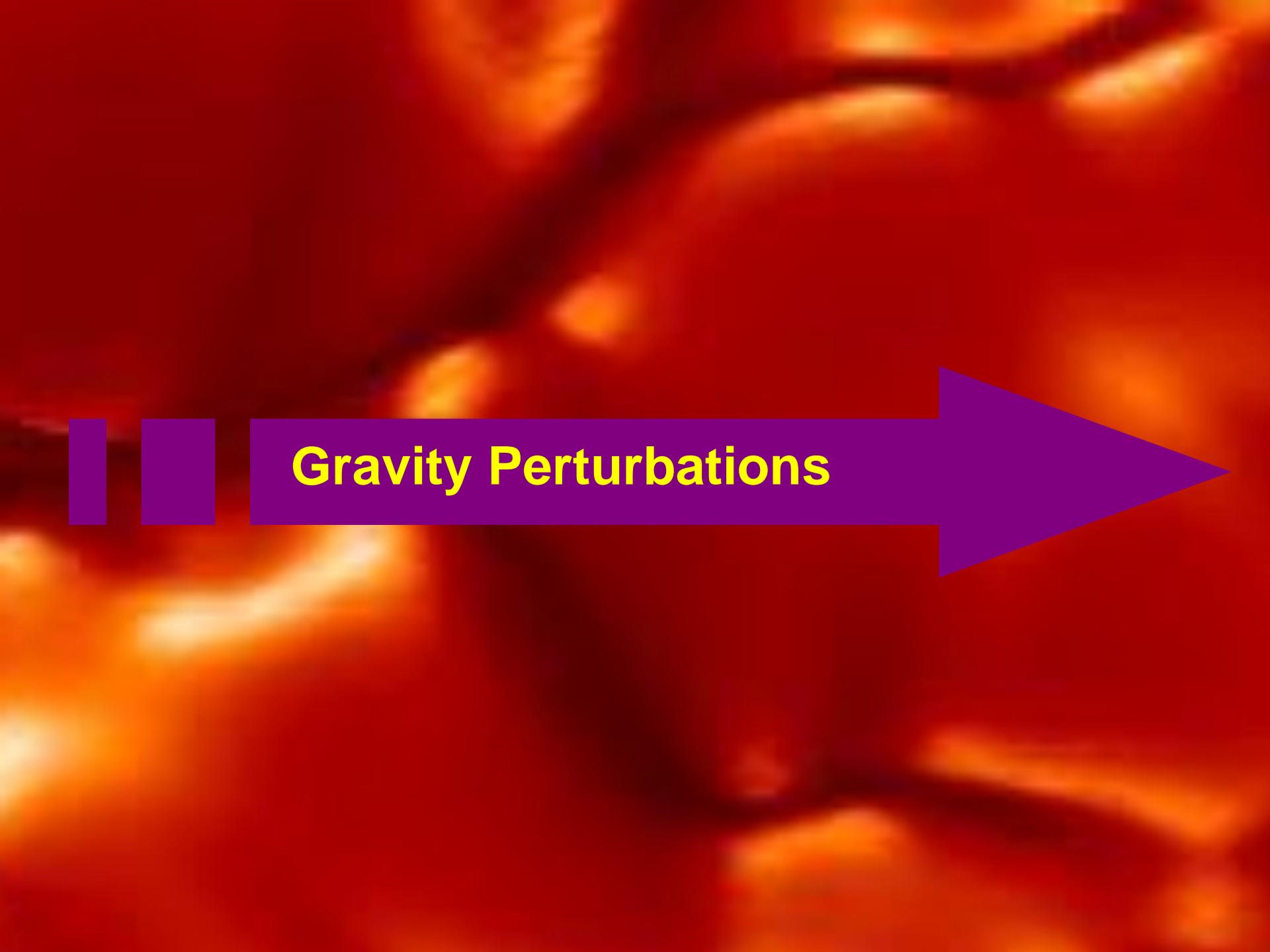
gravitational
origin

full-dome: WMAP-Planck

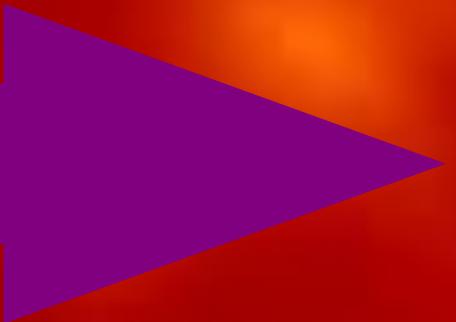


Primordial density fluctuations:

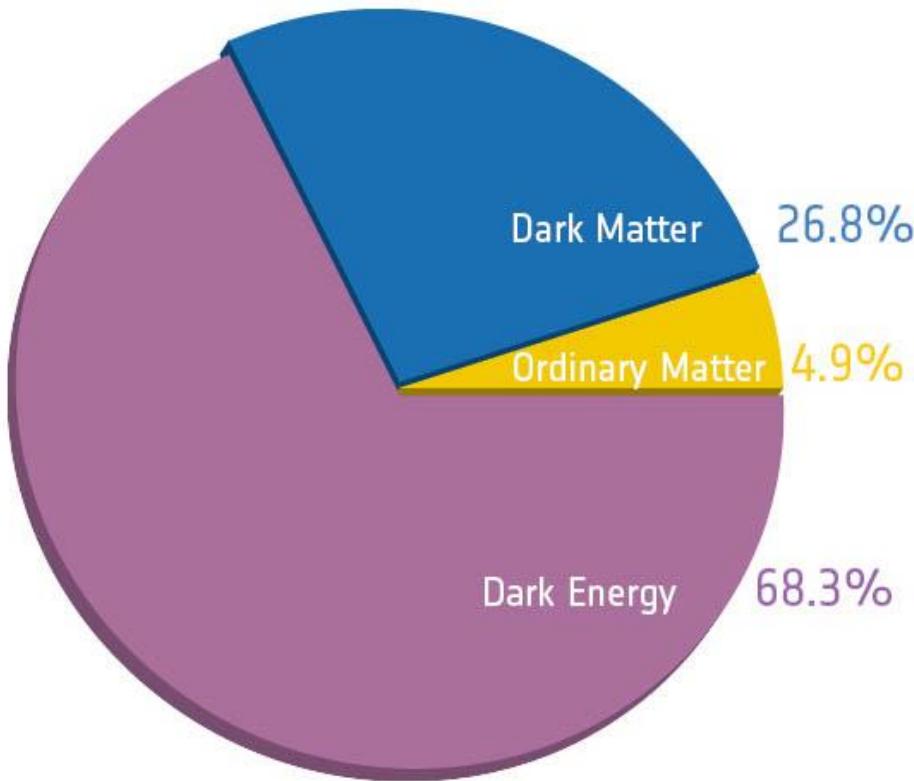
- Origin:
quantum noise generated at inflation
 $t \sim 10^{-36}$ sec.
- visible from 379000 yrs after Big Bang



Gravity Perturbations



Cosmos: the Elements



Composition of the Universe

- **Baryonic (normal) matter** ~ 4.9%
 - atoms of which we ourselves consist
 - stars a mere ~0.20% !
- **Dark Matter** ~ 26.8%
 - gravitationally dominant
 - invisible
 - all structure in the Universe (galaxies, stars, planets, ...) only exist through role of dark matter
 - as yet unknown of which it consists
- **Dark Energy** ~ 68.3%
 - gravitational repulsion
 - dominates expansion (& fate) Universe
 - uniformly distributed over Universe,
 - does not/cannot clump:
less decisive role in structure formation
 - only discovered in 1998 !
- **Straling** ~ 0.001%
 - because of cooling Universe no longer important
 - by far most abundant particle:
 2×10^9 photons per atom

Cosmic Structure Formation

Millennium
Simulation:
LCDM

31.25 Mpc/h

(courtesy:
Virgo/V. Springel).

Dark Matter,
(~ 5.5x more than
baryonic matter)



without: not enough time
to form structure in the
Universe in 13.8 Gyrs

(cosmic web, clusters,
galaxies, stars, ...)

Cosmic Structure Formation

Millennium

Simulation:

LCDM

31.25 Mpc/h

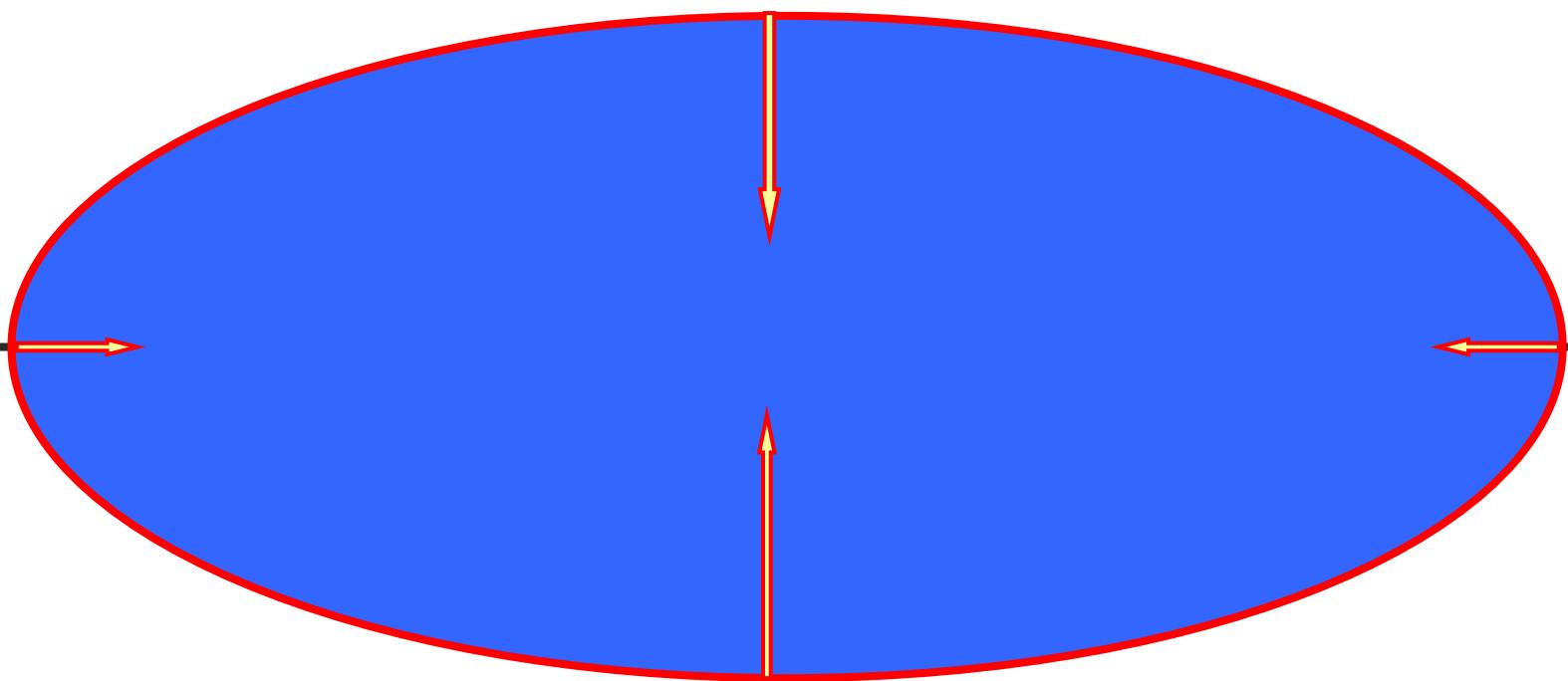
(courtesy:

Virgo/V. Springel).

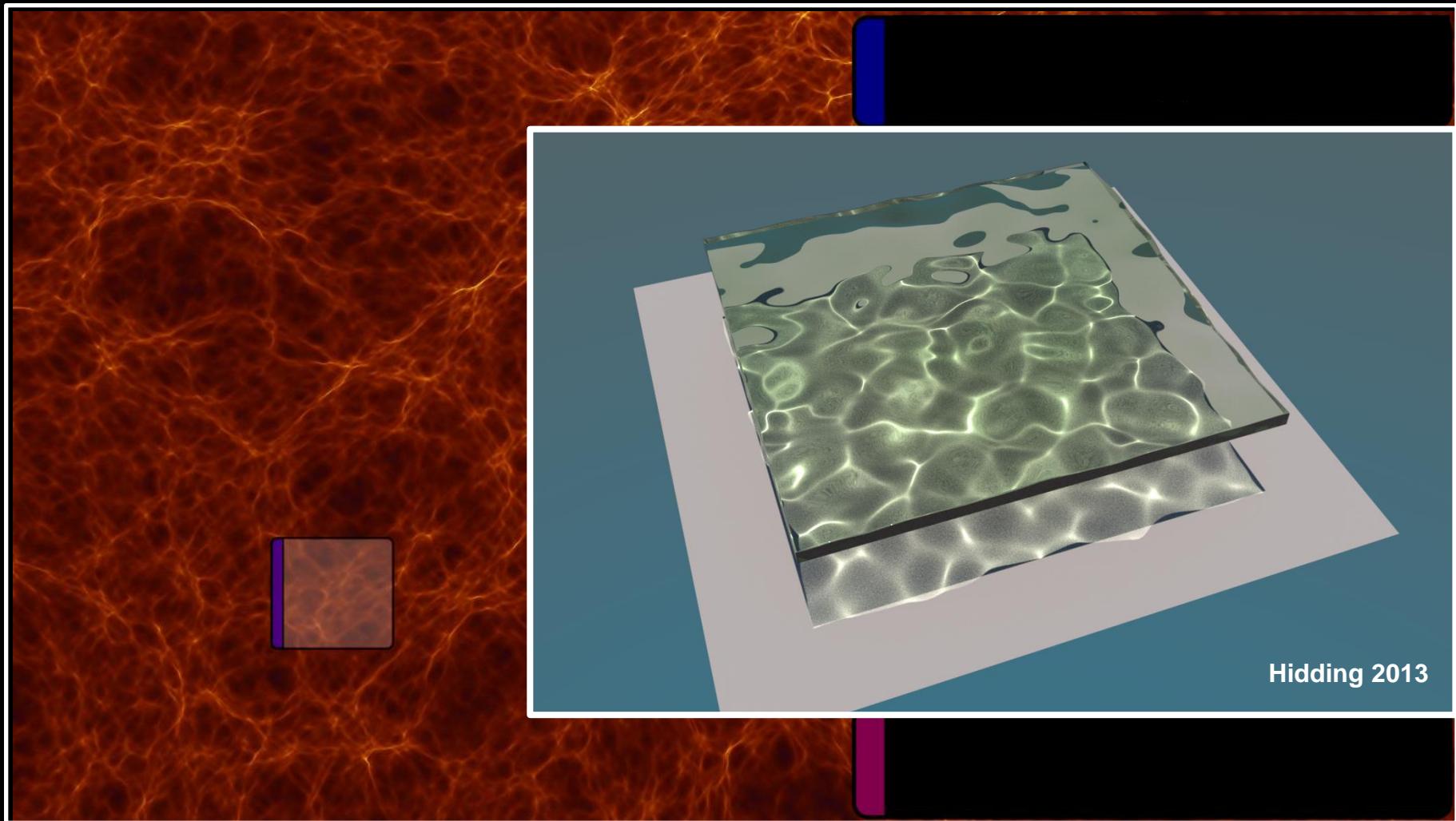
Anisotropic Gravitational Collapse

Amplification

small perturbations in gravity along different directions (tidal forces)

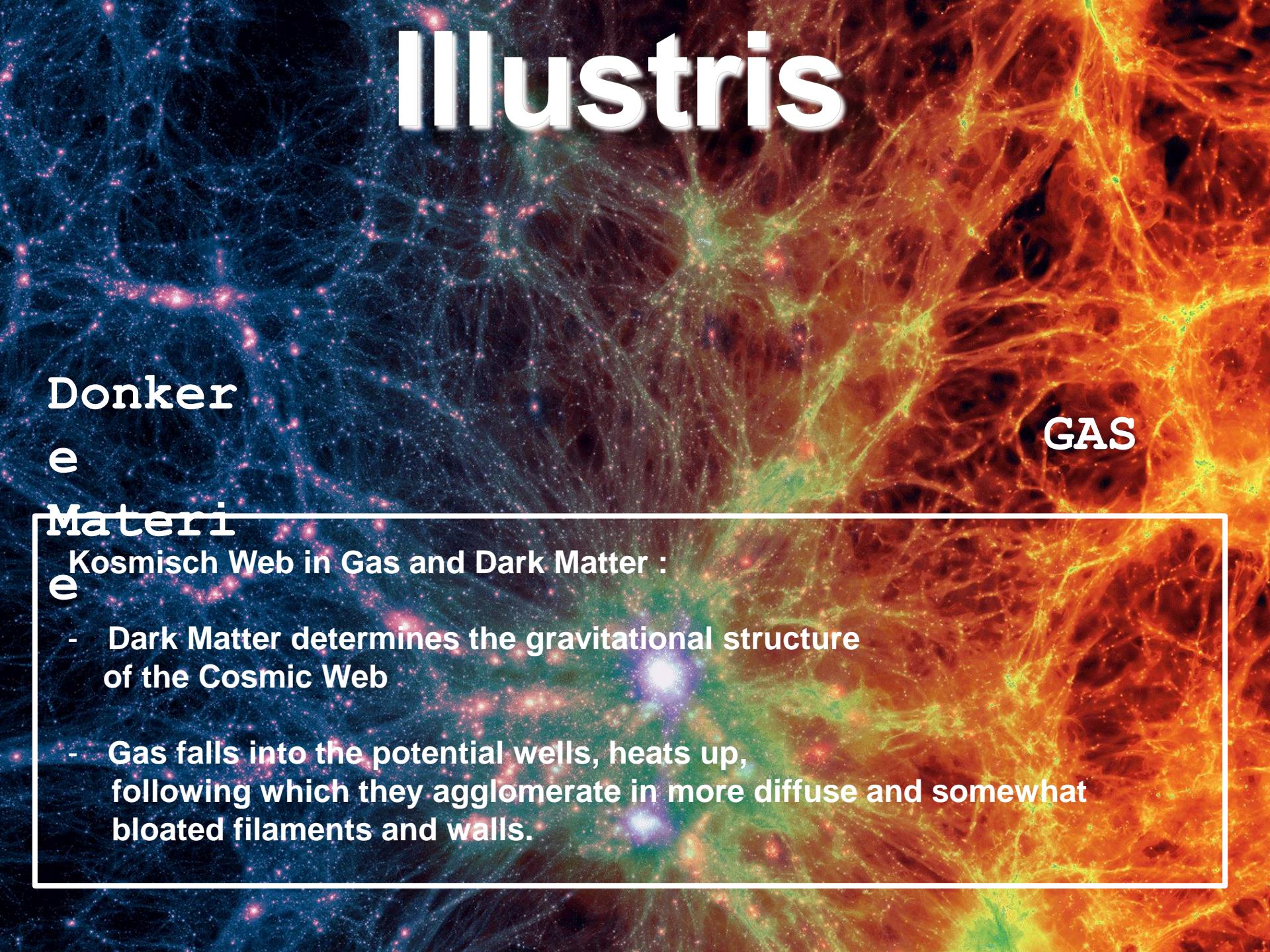


Zel'dovich Approximation



full-dome: Zeldovich

Illustris

The background image shows a simulation of the Cosmic Web. It features a complex network of filaments and voids. The filaments are primarily colored in shades of orange and red, representing gas density. The voids between these filaments are mostly blue and green, representing lower density regions. Small white dots scattered throughout represent individual galaxies.

Donker

e

Materi

e Kosmisch Web in Gas and Dark Matter :

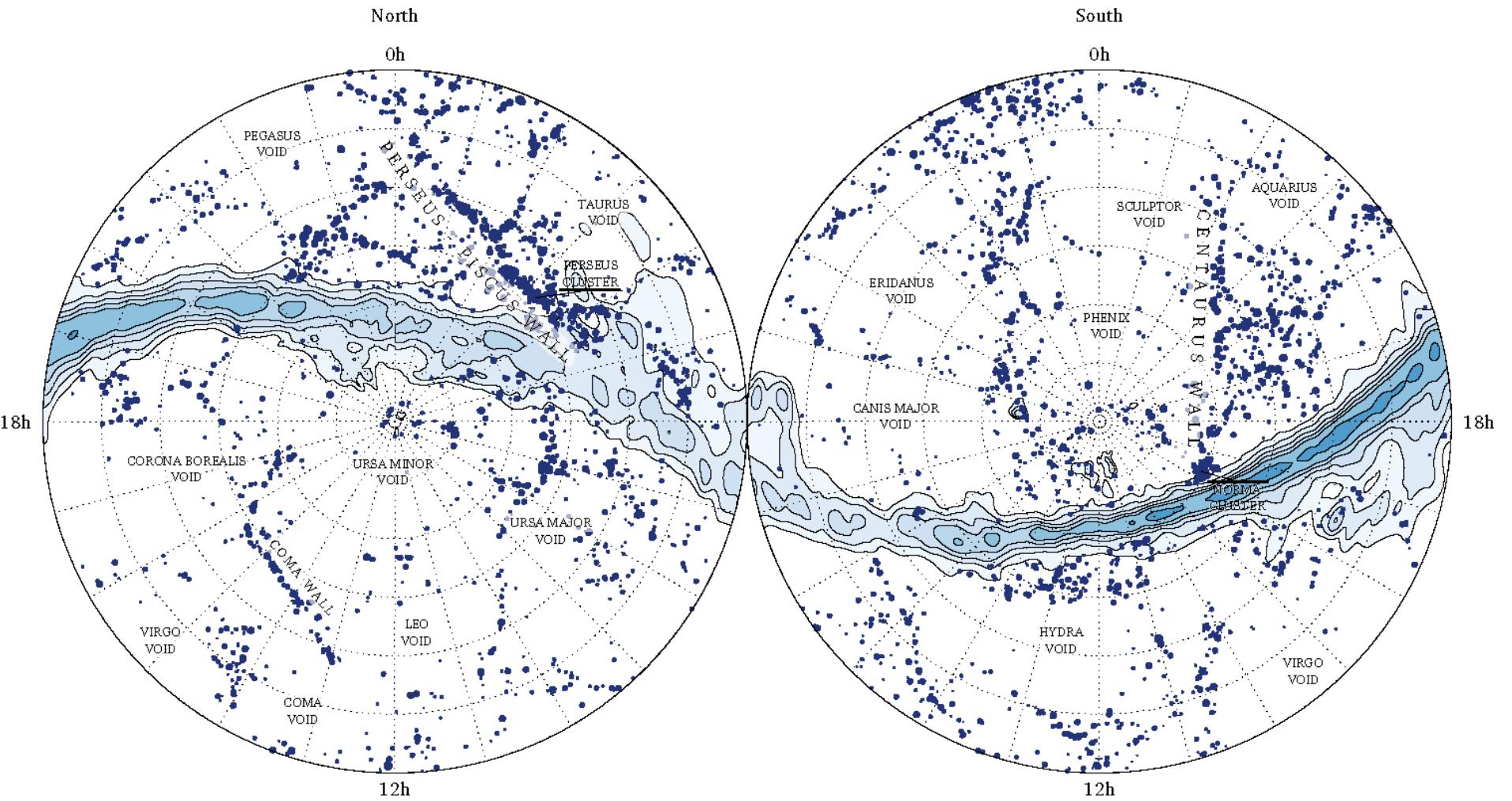
- Dark Matter determines the gravitational structure of the Cosmic Web
- Gas falls into the potential wells, heats up, following which they agglomerate in more diffuse and somewhat bloated filaments and walls.

GAS

full-dome: Illustris

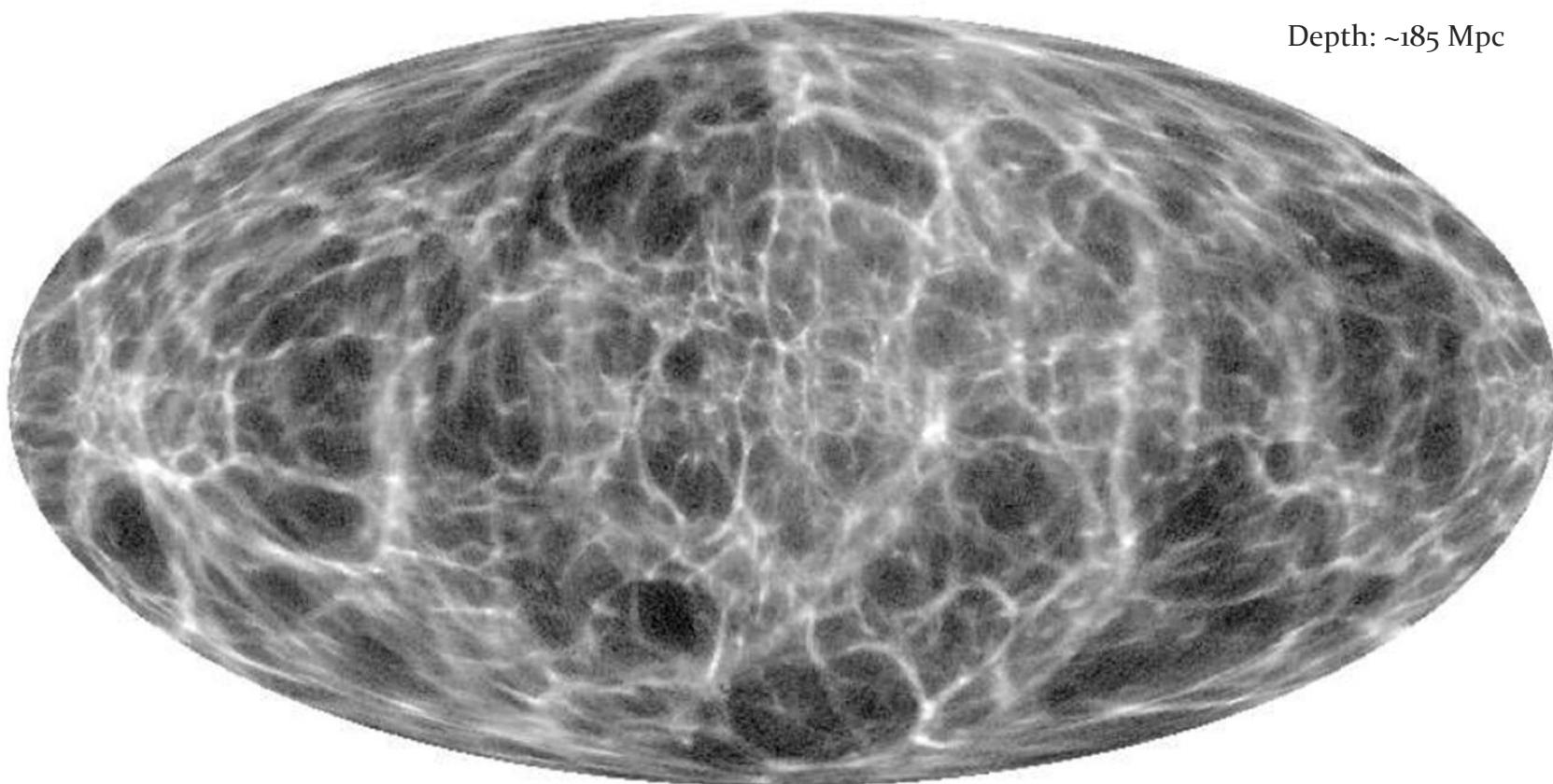
Cosmic Web: Local Universe

Galaxies up to 50 Mpc ...



2MRS survey sky map: Hidding 2015

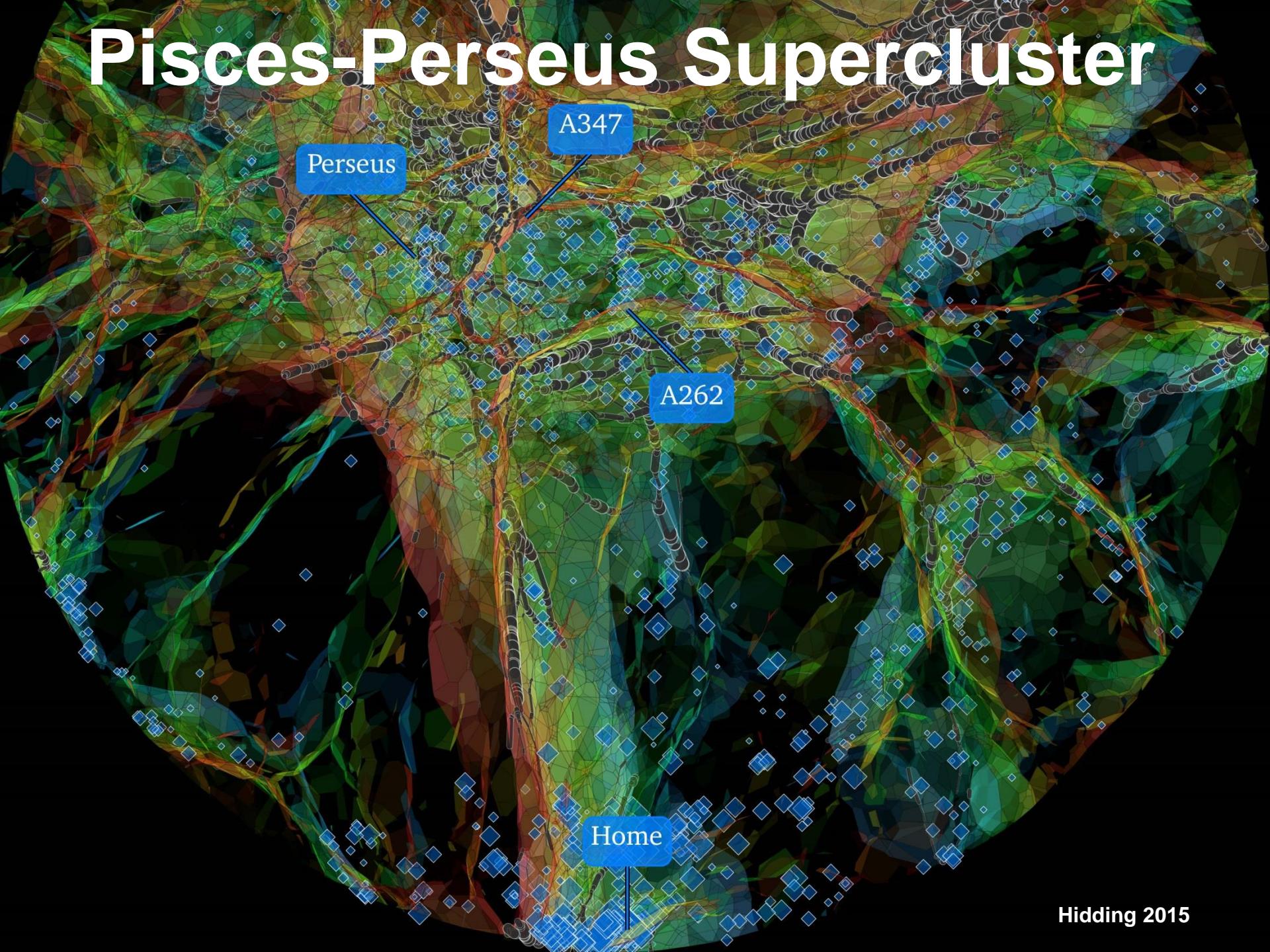
Caught within our Web ...



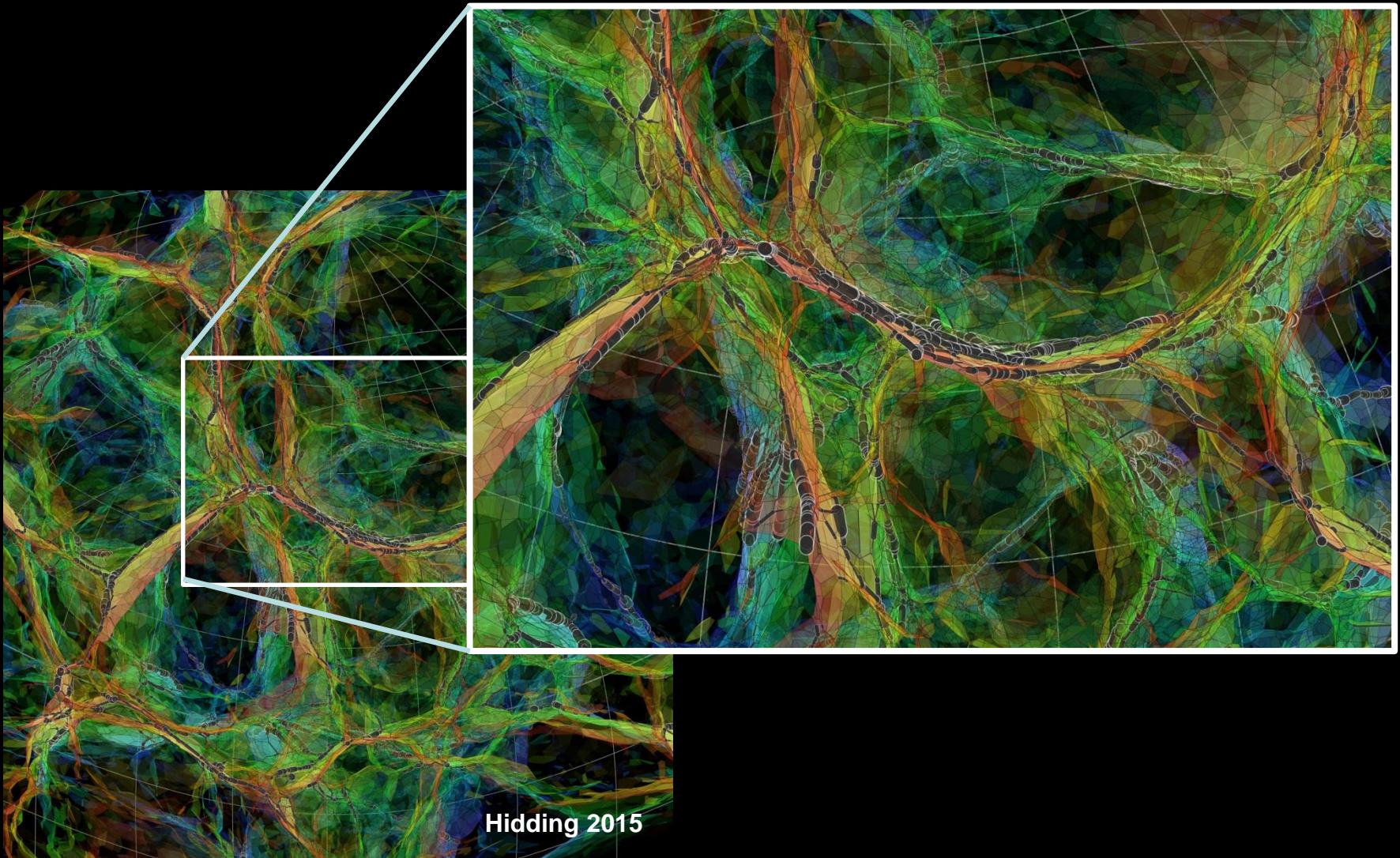
Most detailed
Reconstruction
Local Cosmic Web:

Depth: ~185 Mpc

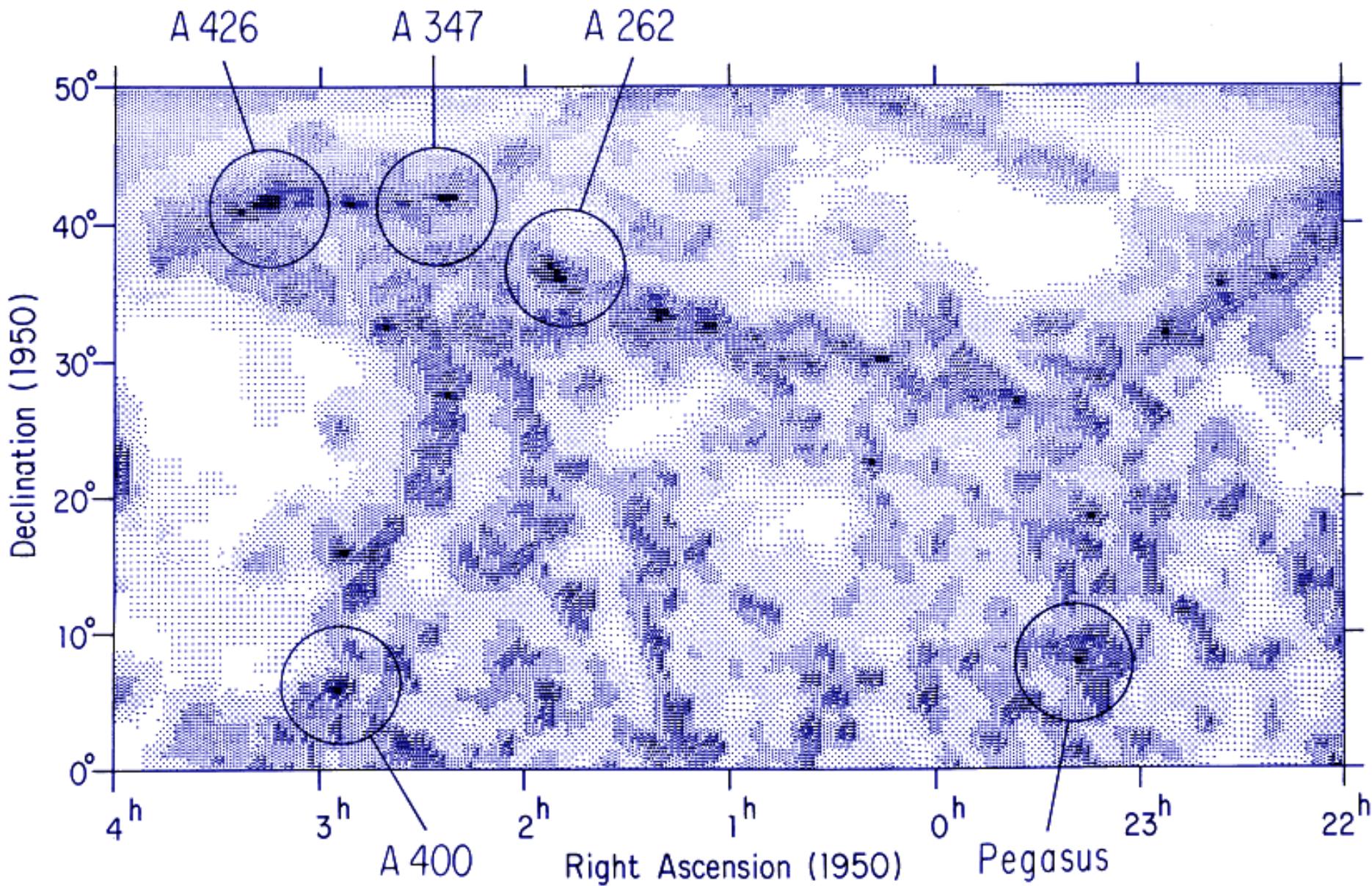
Pisces-Perseus Supercluster



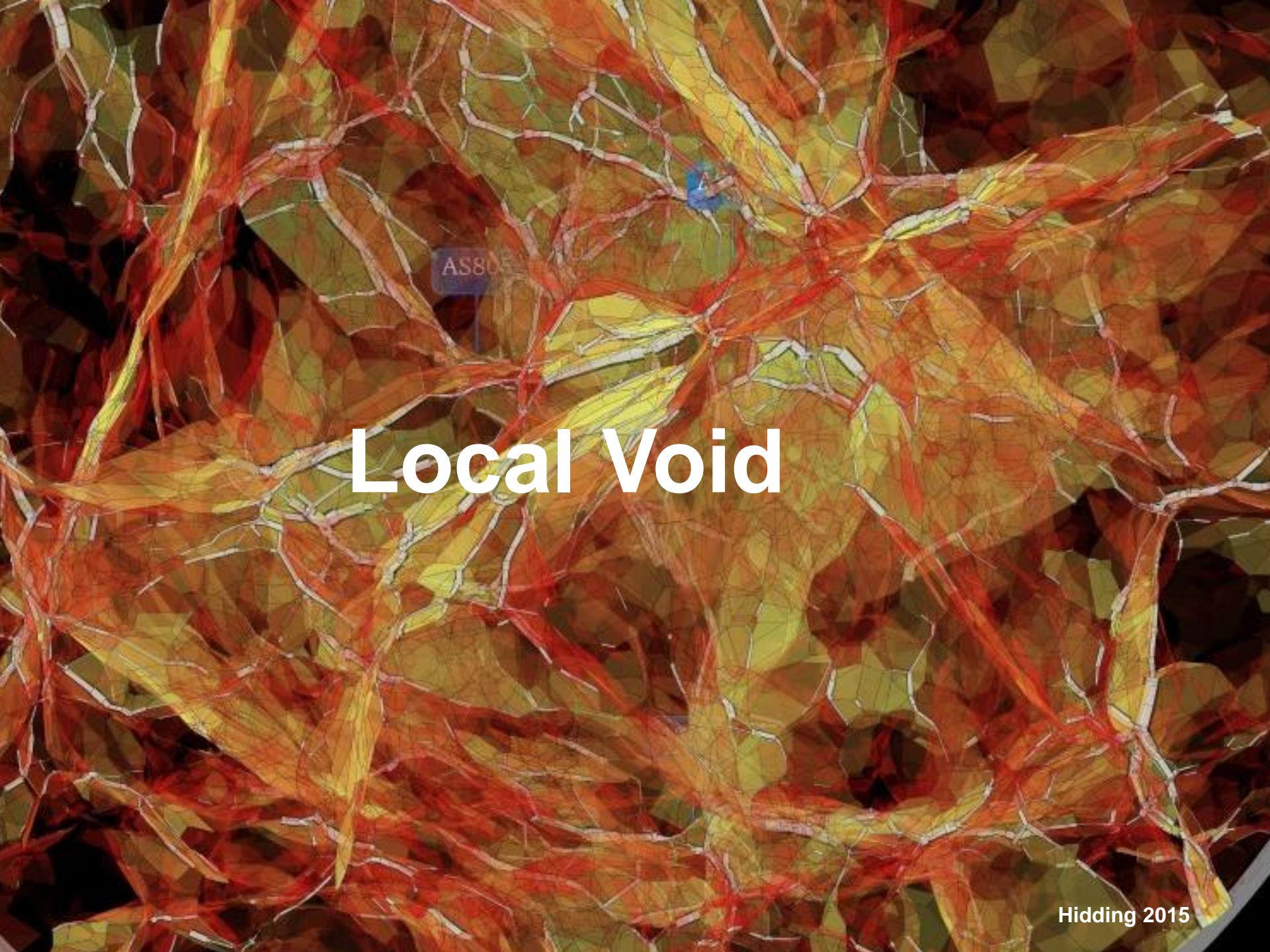
Pisces-Perseus Supercluster



Pisces-Perseus Supercluster



full-dome: Johan-Pisces Perseus



AS80

Local Void

Local Void

- Voids are prominent aspects of the Cosmic Web, instrumental in spatial organization of the Megaparsec Universe.

Sheth & van de Weygaert 2004

Hidding, van de Weygaert, Kitaura & Hess 2015

- Voids contain significant amount of information on global cosmological parameters:
 - void outflow: dark matter
 - void shapes: dark energy
 - supervoids: existence

Bos, van de Weygaert, Dolag & Pettorino 2012

- Voids are pristine low-density regions, ideal for studying galaxy formation and the effects of cosmic environment on the formation of galaxies.

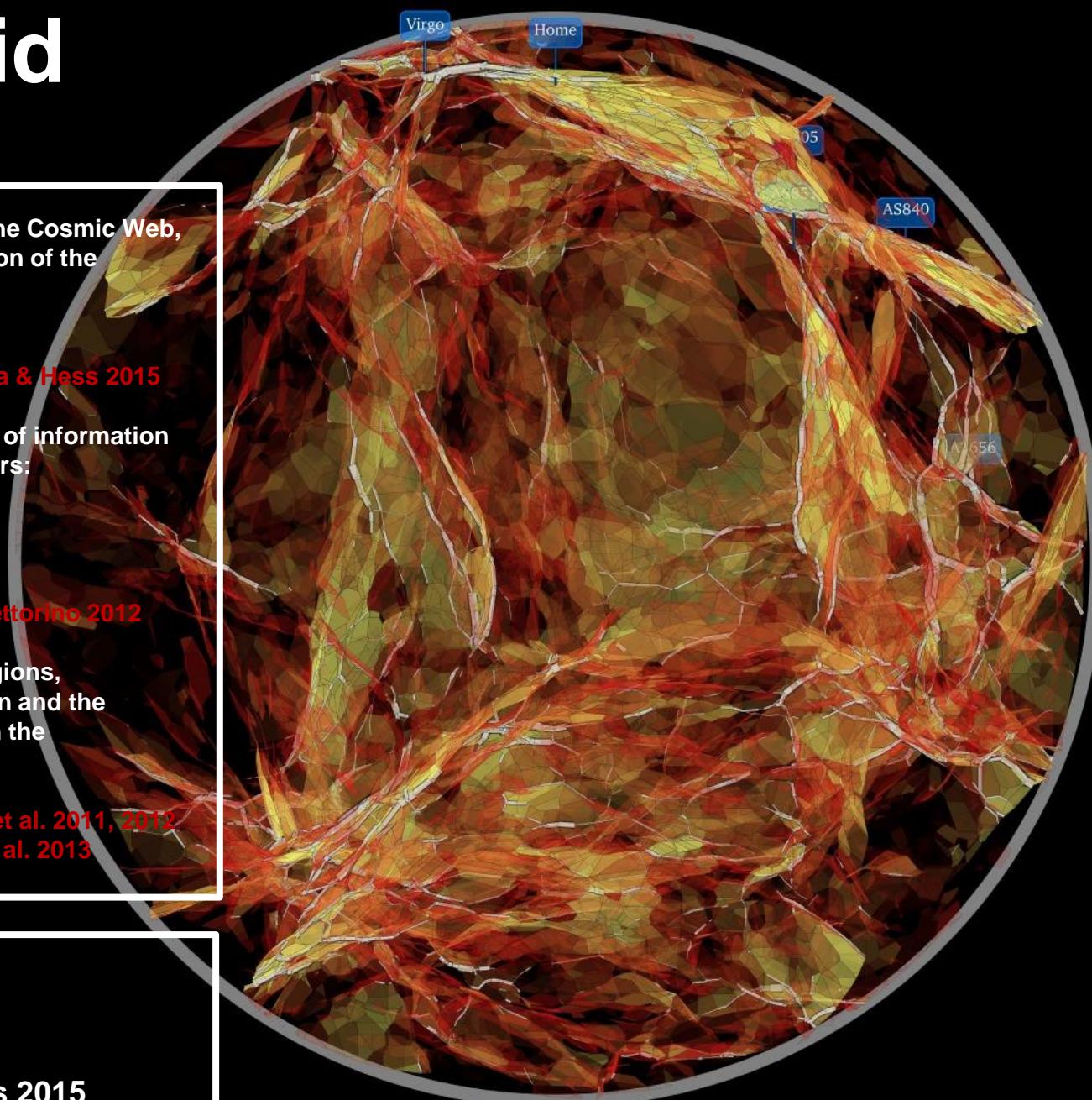
Void Galaxy Survey: Kreckel et al. 2011, 2012

Beygu et al. 2013

Lokale Void

Reconstructie:

Hidding, vdW, Kitaura & Hess 2015



full-dome: Johan-Local Void



Thank you for your attention

Are there any questions ?

image courtesy:
Aragon-Calvo, Subbarao & Szalay

full-dome: SDSS-2 (flythrough, music)